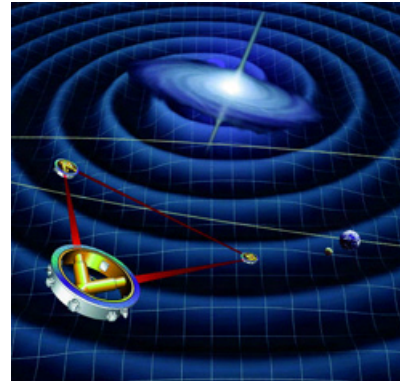


Effects of Ultraviolet Irradiation on LIGO Mirror Coating



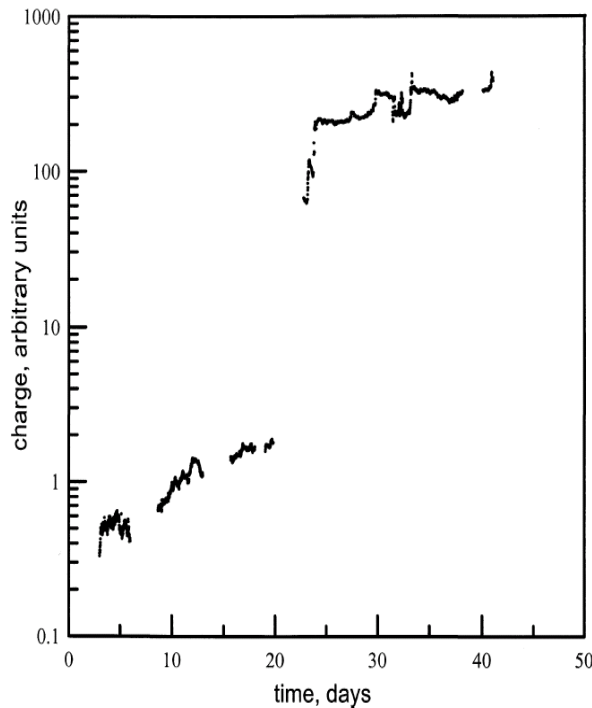
Ke-Xun Sun, Nick Leindecker, Ashot Markosyan,
Sasha Buchman, Roger Route, Marty Fejer, Robert Byer,
Helena Armandula, Dennis Ugolini, Gregg Harry

Workshop on Optical Coatings in Precision Measurements
California Institute of Technology, Pasadena CA
March 20-21, 2008



LISA, LIGO and Adv LIGO Charging Problems

Time dependence of electric charge on the test mass



| Charging | LISA | LIGO | Adv LIGO |
|--|------------------------------------|-----------------------------|-----------------------------|
| Mechanism | Space Weather Caging Separation | Cosmic Ray Triboelectric | Cosmic Ray Triboelectric |
| Charging Rate | 10^{-11} C/day (max) | 10^{-7} C/day | 10^{-6} C/day |
| Test Mass | 2 kg | 10 kg | 40 kg |
| Displacement Sensitivity | 10^{-11} m | 10^{-18} m | 10^{-20} m |
| Frequency | 0.03 mHz~1 Hz (1 mHz) | 10Hz ~ 1kHz (100Hz) | 30~2 kHz (30Hz) |
| Figure of Problem (C/M ω^2)/ Δx | 13 | 25 | 6.8×10^4 |

Moscow Data

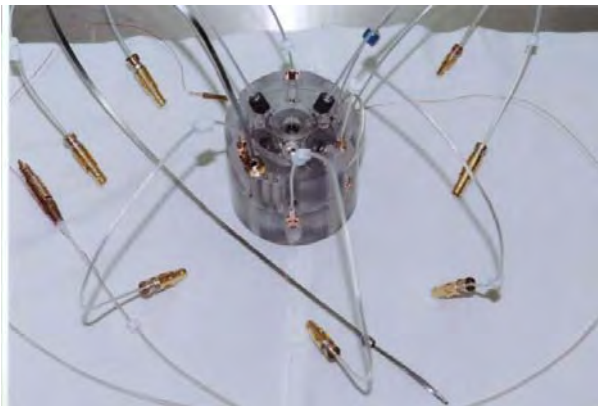
Advanced LIGO may see more charging problems instead of less



Charge Management In Precision Experiments

GP-B charge management

- **Critical to GP-B mission success**
 - > **Initial gyro lifting-off**
 - > **Continuous charge management during science measurement**



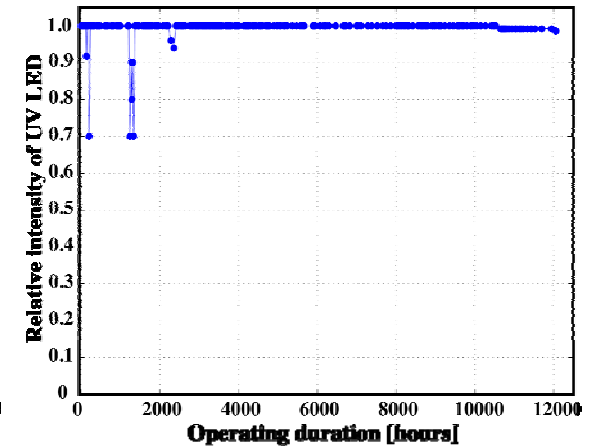
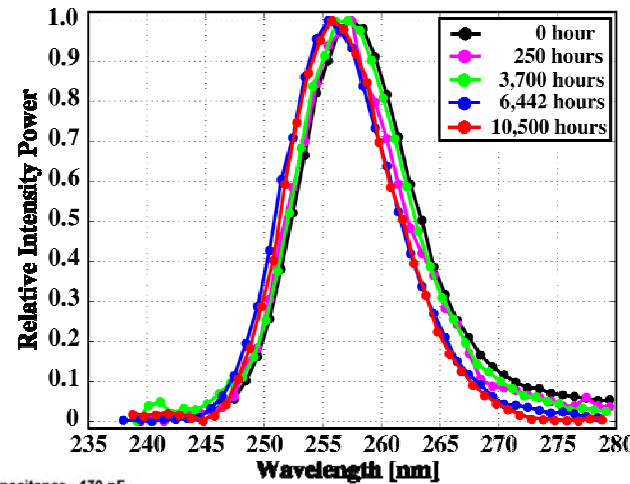
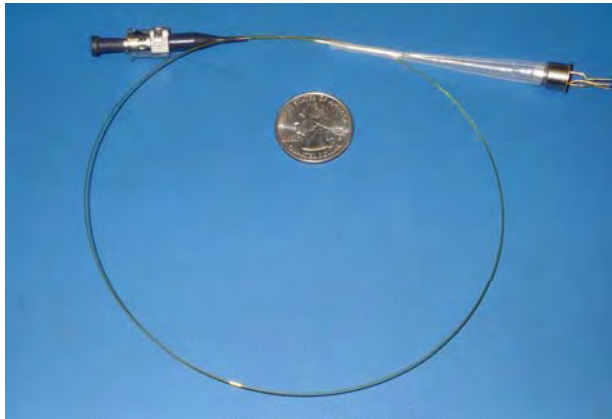
LIGO UV charge management budget

- $Q_c \sim 10^{-7} \text{ C/m}^2$ commonly cited
- Charging rate $Q_c \sim 10^{-7} \text{ C/day}$
- $N_e \sim 10^{12}$ electrons/day
- Photoelectric “Q. E.”: $\eta \sim 10^{-6}$
- UV photons required: $N = 10^{17}$
- $P_{UV} = Nhc/\lambda T = 8.9 \times 10^{-6} \text{ W}$
- $P_{UV} \sim 10 \text{ } \mu\text{W}$ (average power over a day)
- **Dynamic Range ~ 100**
- $P_{UV} \sim 1 \text{ mW}$ (Peak power)

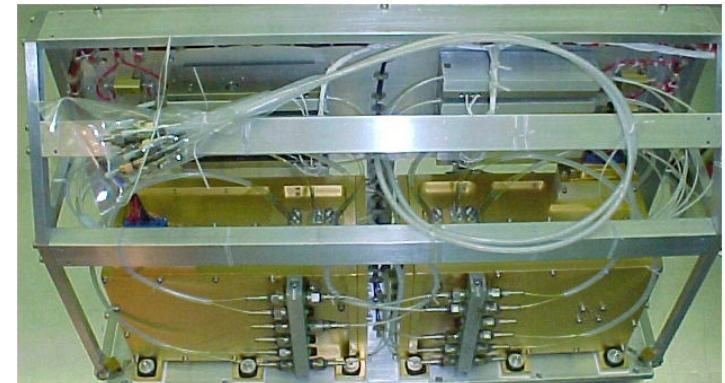
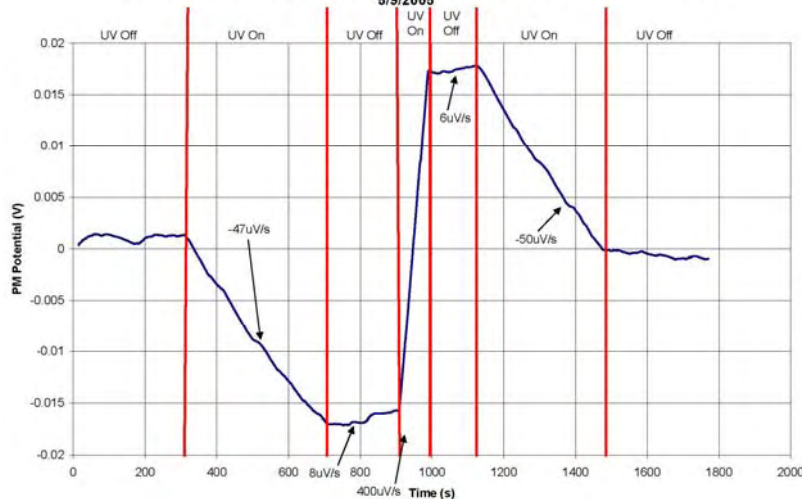
Stanford experiences in charge management: GP-B, ST7, LISA (MGRS)



UV Light Source UV LED & Gas Lamps



AC Charge transfer 10KHz UVLED 1.2 mA 50% duty cycle System Capacitance ~170 pF, 5/9/2005

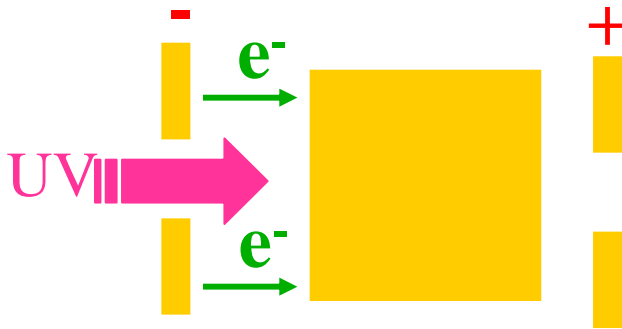


Ke-Xun Sun, Brett Allard, Scott Williams, Sasha Buchman, and Robert. L. Byer, "LED Deep UV Source for Charge Management for Gravitational Reference Sensors," presented at Amaldi 6 Conferences on Gravitational Waves, June 2005, Okinawa, Japan, Class. Quantum Grav. **23** (2006) S141-S150

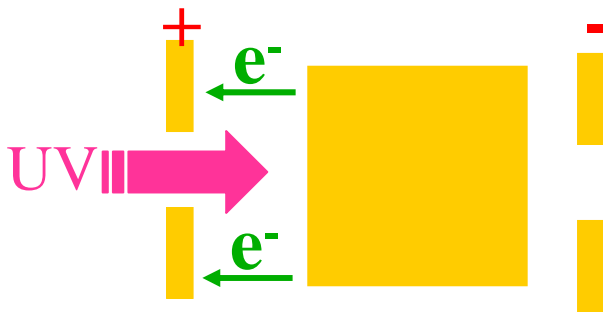
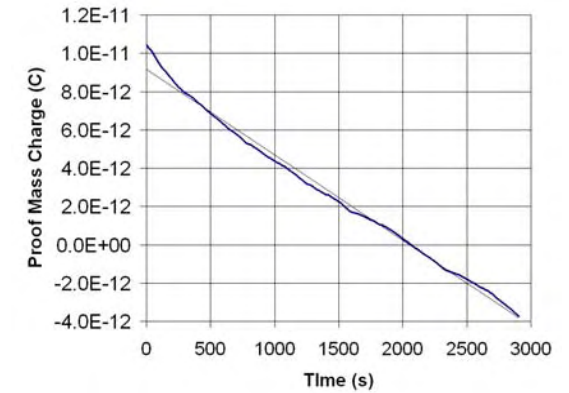
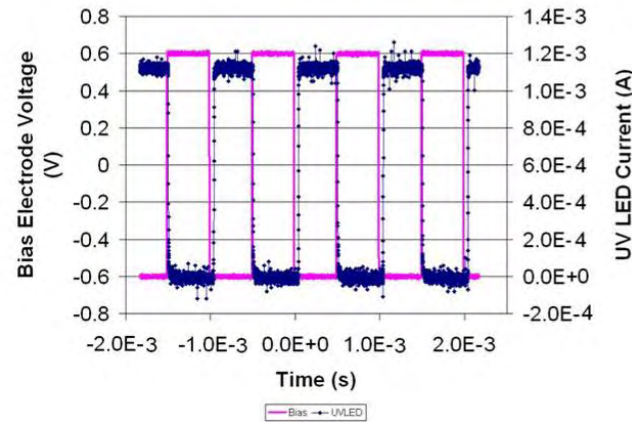


AC Charge Management

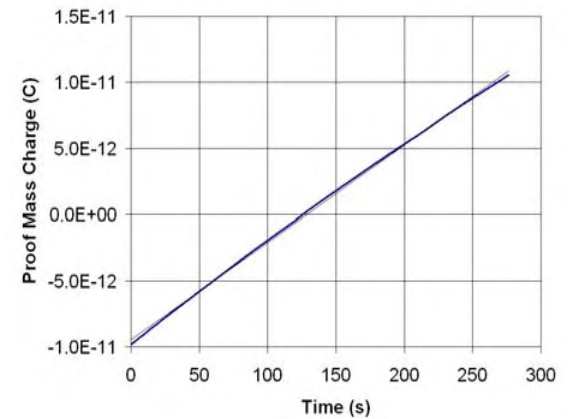
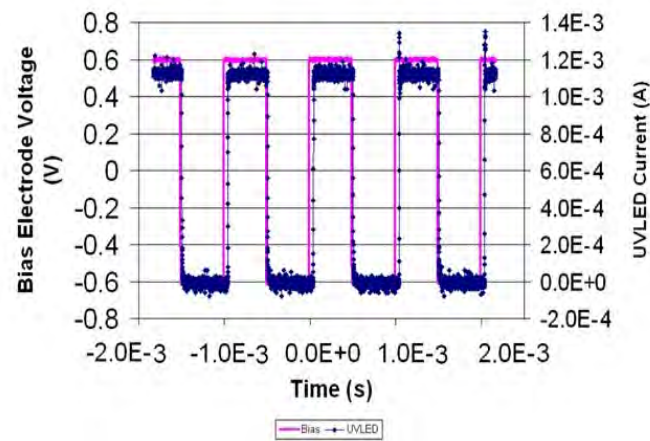
UV LED and bias voltage modulated at 1 kHz



May 6, 2005 Negative Charge Transfer Phasing



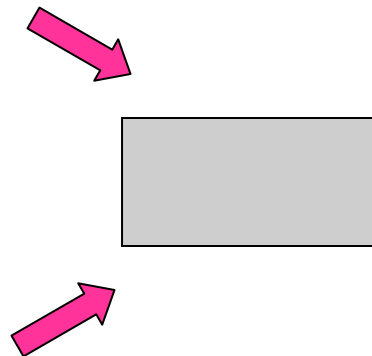
May 6, 2005 Positive Charge Transfer Phase Configuration



UV Illumination Configurations

• Direct illumination

- UV mercury lamp is routinely used for attachment removal
- UV LED has sufficient power for cw direct illumination
- Works for removing charges
 - Metal: GPB, ST7, LISA
 - Dielectric: Glasgow, Trinity, GEO, Moscow
- UV irradiation may cause problem to coating



• Indirect illumination

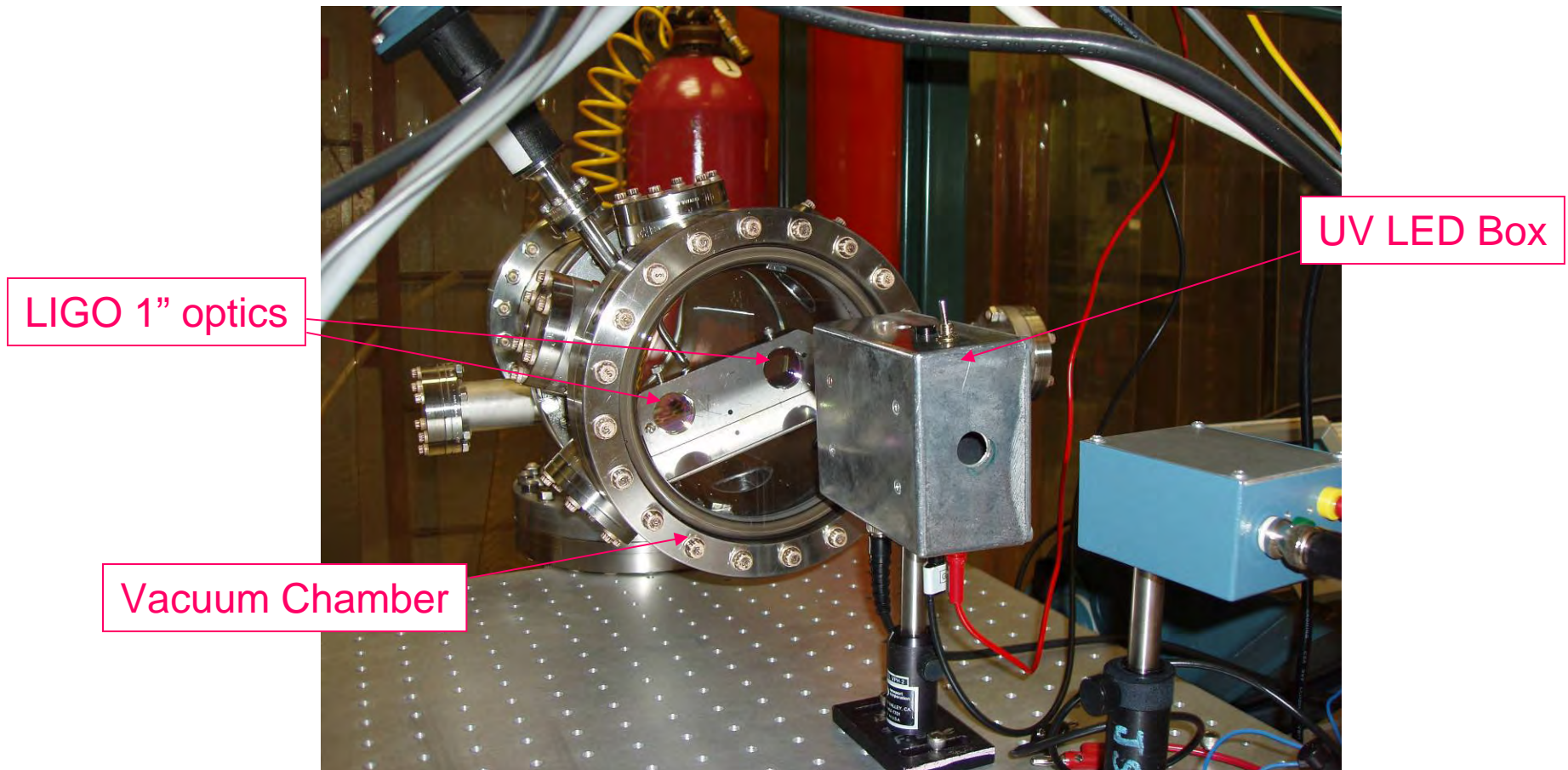
- Au coating on non-critical portions of test mass and suspension structure (P. Willem: 0.1 μm Au coating on barrel, opaque to UV)
- Photoelectric effect on Au surface has been utilized in GP-B, ST-7, LISA
- Establish electric field to herd the charges to Au coating
- UV photoelectric effect continues to remove Charges
- Higher throughput in charge control
- No UV illumination on dielectric



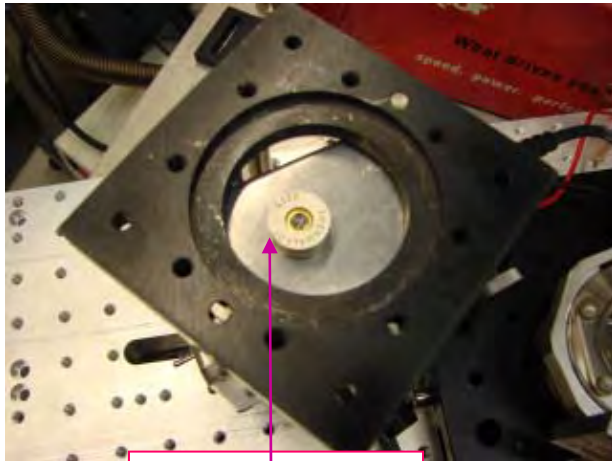
- **Three samples irradiated with UV light**
 - **REO #2: Ta₂O₃/SiO₂ alternative layers**
 - **Adv LIGO (LMA: 14 TaO:Ta₂O₃/SiO₂)**
 - **REO #1: Ta₂O₃/SiO₂ alternative layers**
 - > **UV LED**
 - > **Xeon Lamp**
 - > **Heat/Cool cycles**
- **Absorption loss measured with PCI**
 - **Alternate measurement with UV exposure**
 - **During measurement the samples are controlled under vacuum**



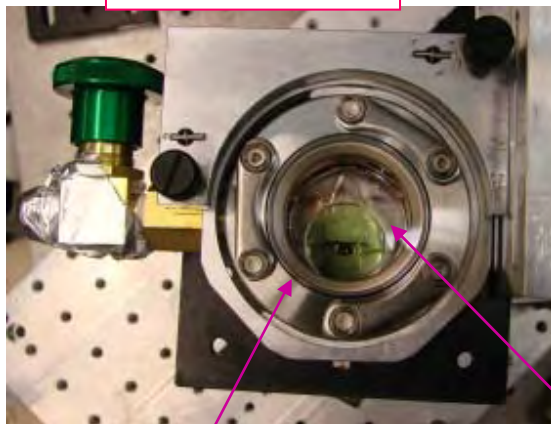
The Initial UV Illumination Geometry



More Localized UV Exposure Geometry



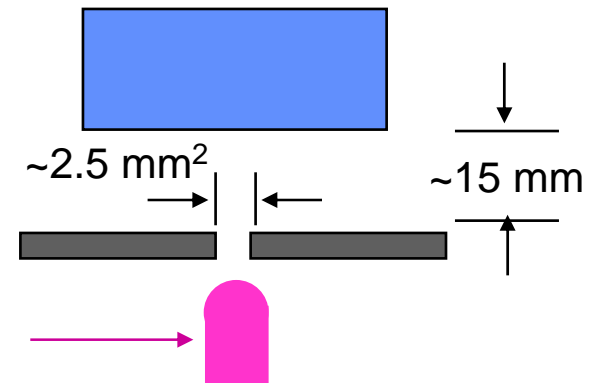
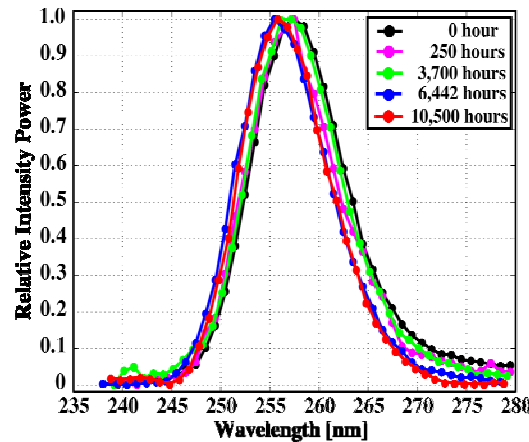
UV LED with a Ball Lens



Small aperture localizes irradiation

Fused silica window with 93% UV Transmission

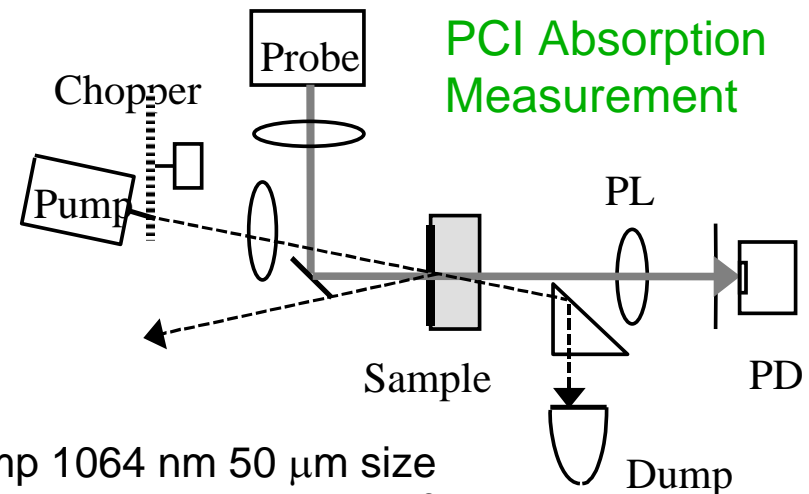
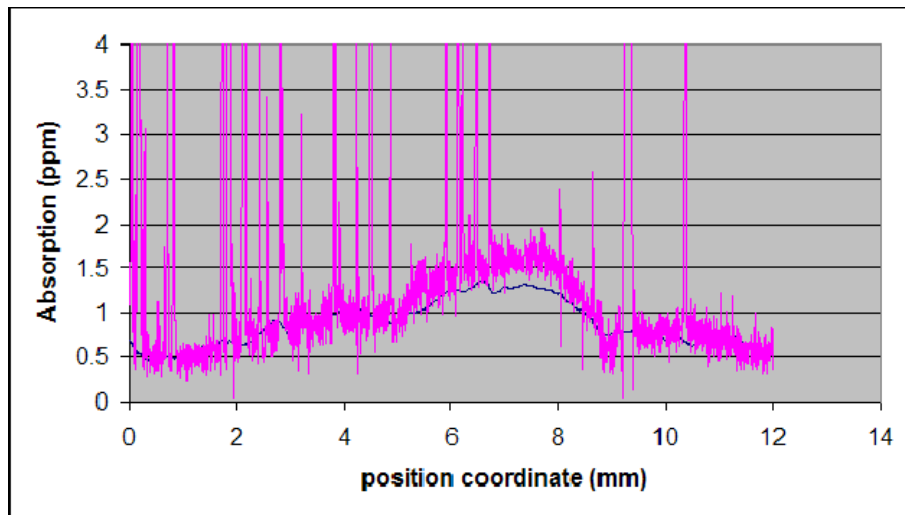
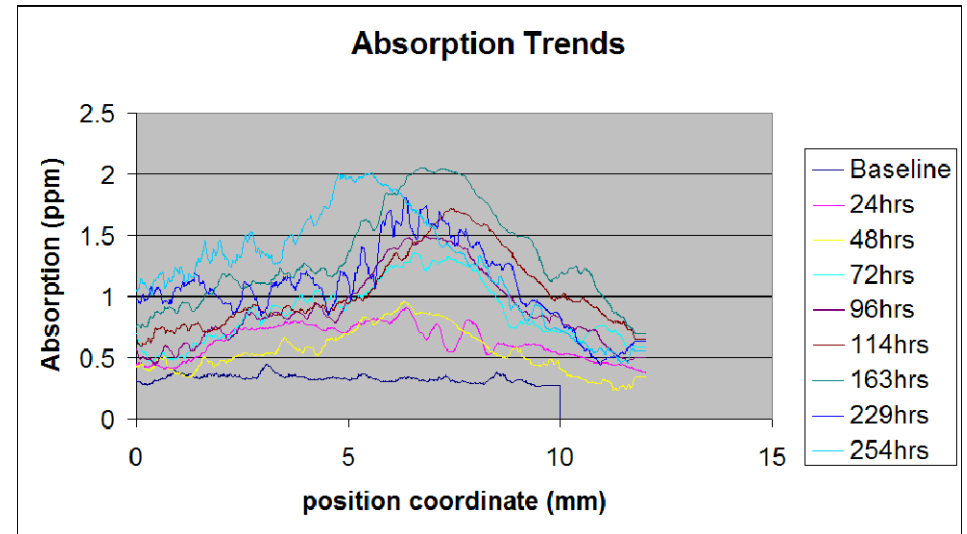
Initial LIGO recipe coated by REO (2 pcs)
HR/AR OX936/OX937
Ta₂O₅/SiO₂



Irradiation by UV LED at 255 nm
Operated with average power 3.3 μ W
Irradiation intensity $\sim 15 \mu$ W/cm²
Irradiation for 254 hours
Total UV energy deposition ~ 16 J/cm²
(> 1 J/cm² recommended by Dennis Ugolini)

UV Exposure and Loss Measurement

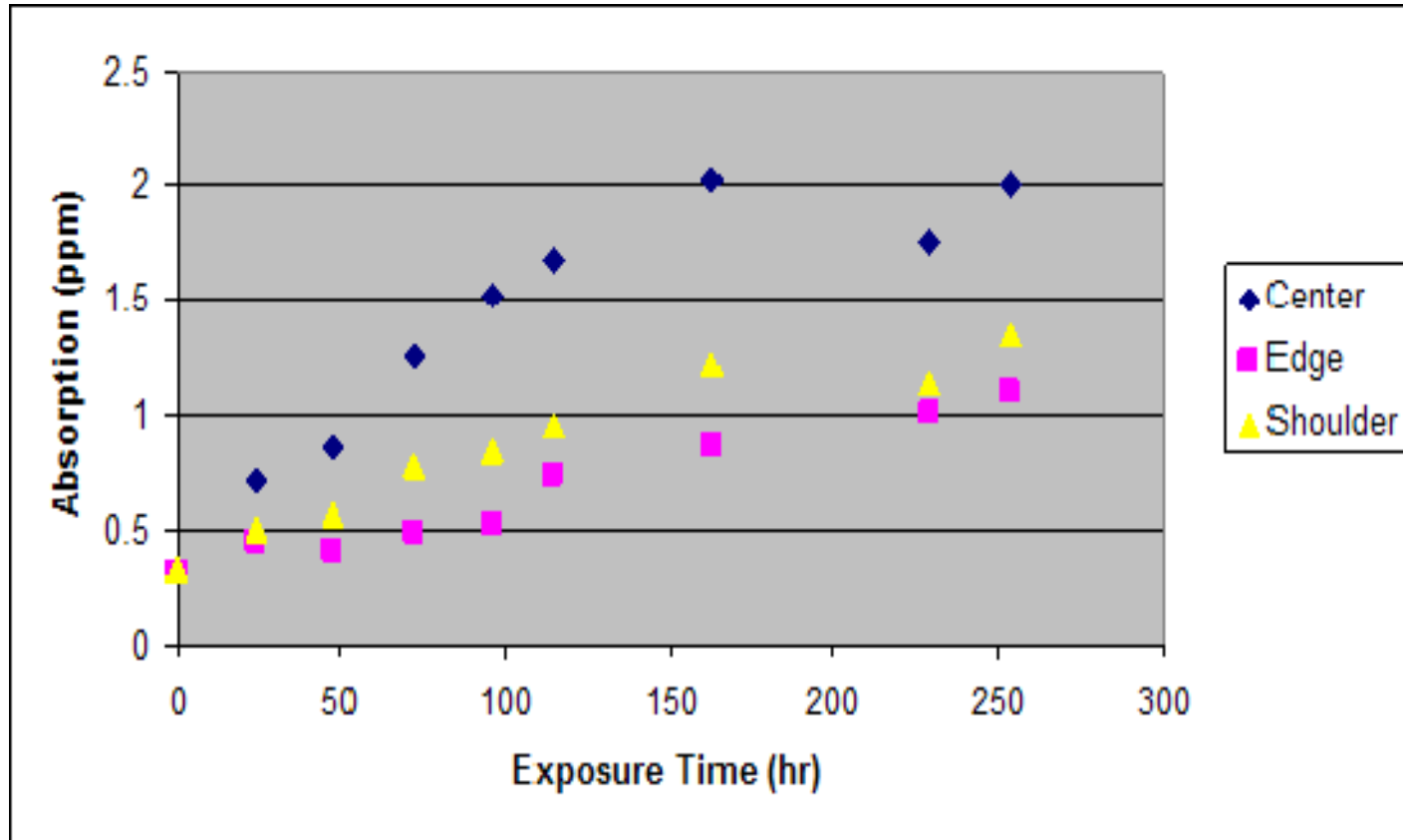
- Alternate UV exposure and loss measurement
- UV exposure was conducted in vacuum 10^{-7} torr
- UV effect after in air exposure may be more recoverable
- Total UV exposure 16 J/cm^2 (254 hours)
- Equivalent time span for LIGO charge management: 4~40 months



Pump 1064 nm 50 μm size
Power density 460 kW/cm^2



UV Effect on Initial LIGO Sample (REO #2)

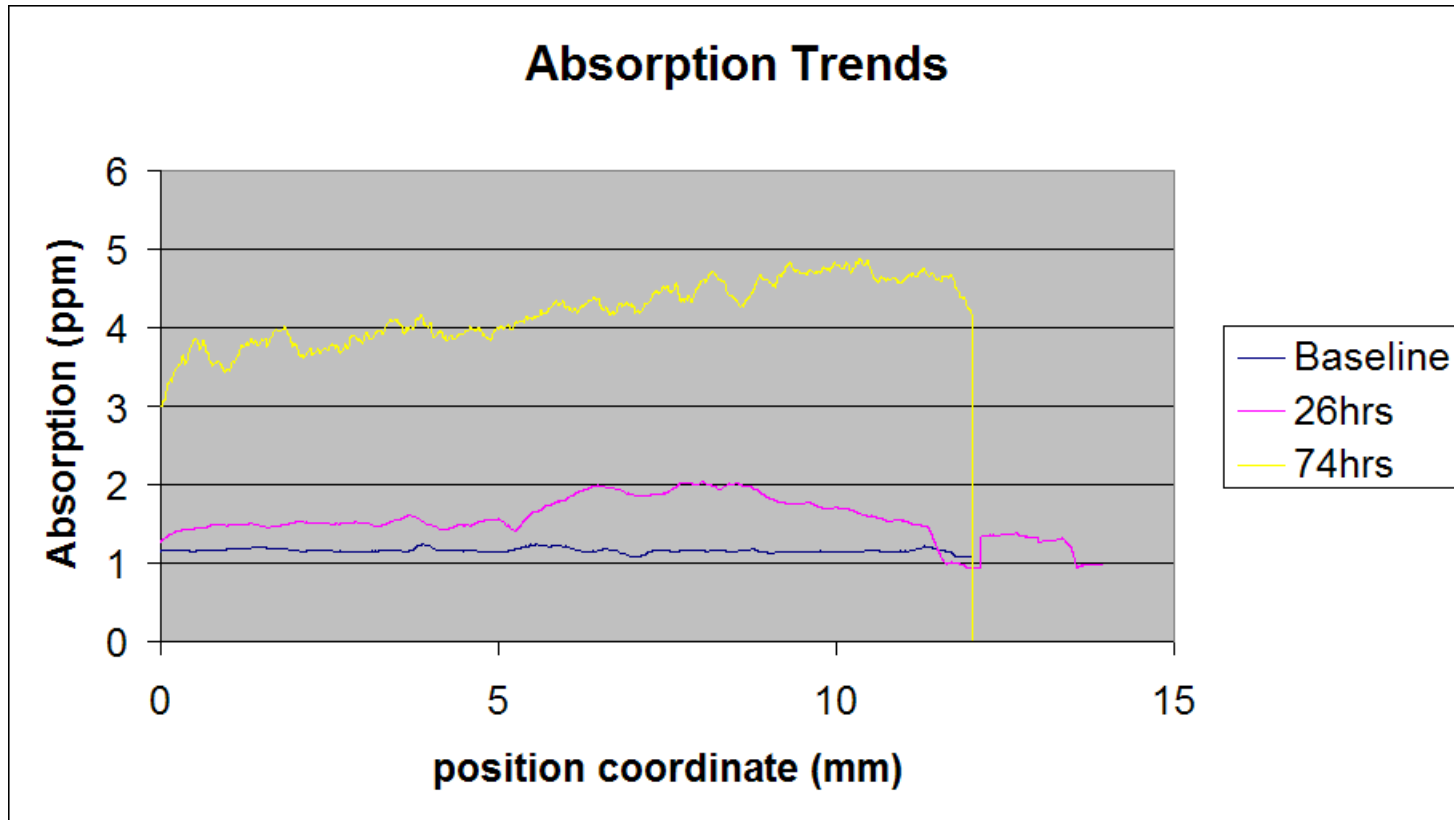




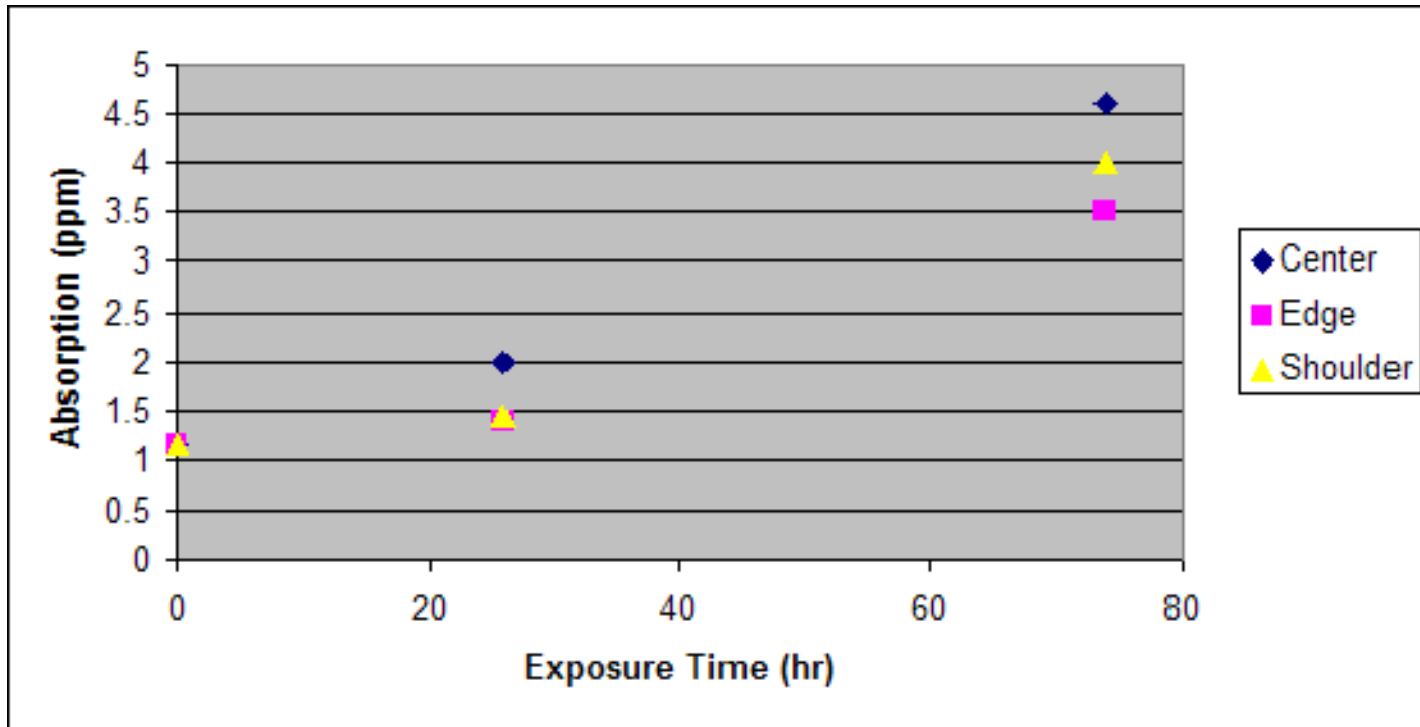
UV Irradiation: Adv LIGO Sample: Spatial Profile



TaO:Ta₂O₃/SiO₂ 30 Layers

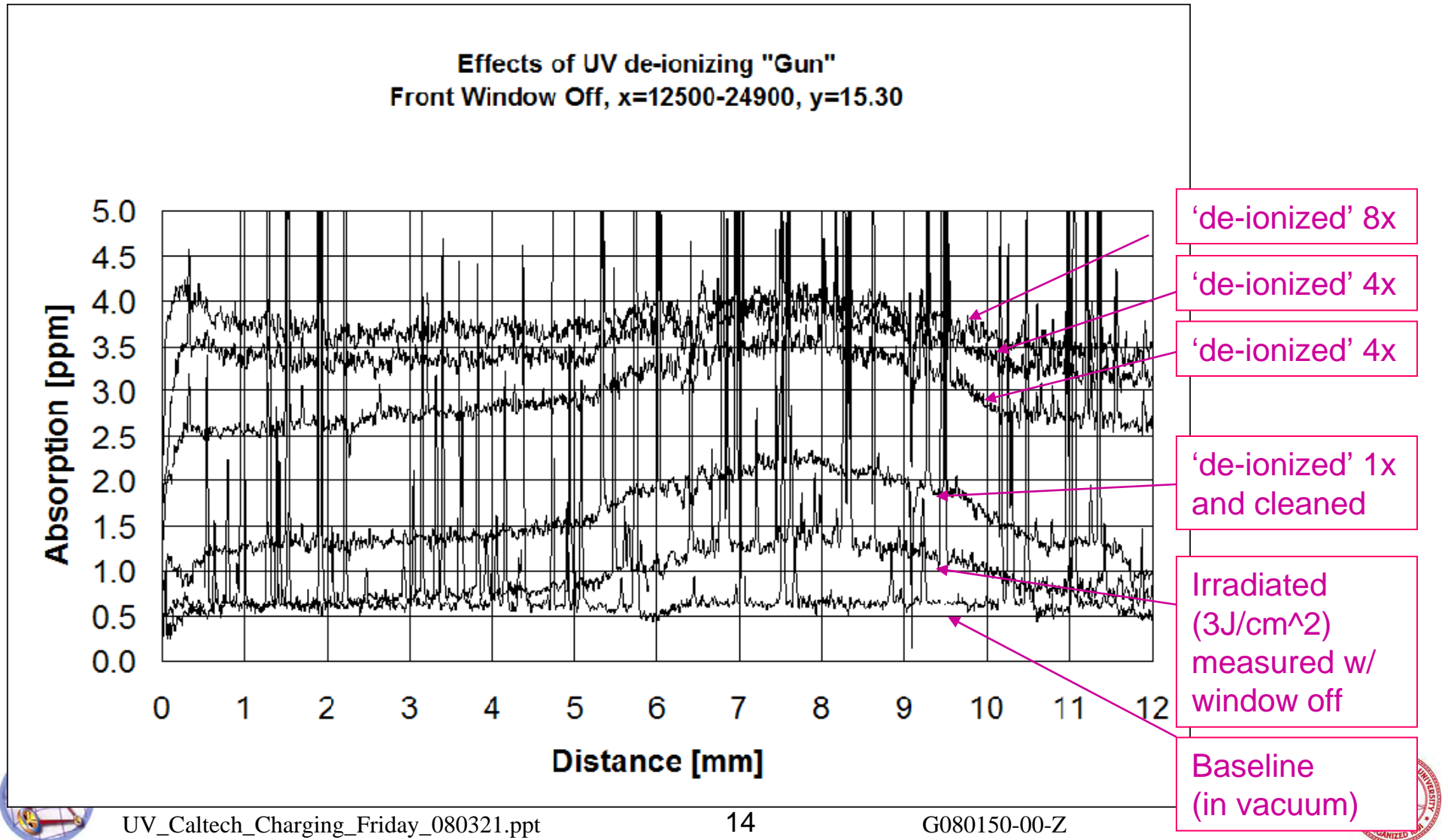


UV Irradiation: Adv LIGO Sample: Temparo Profile

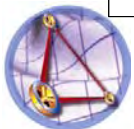
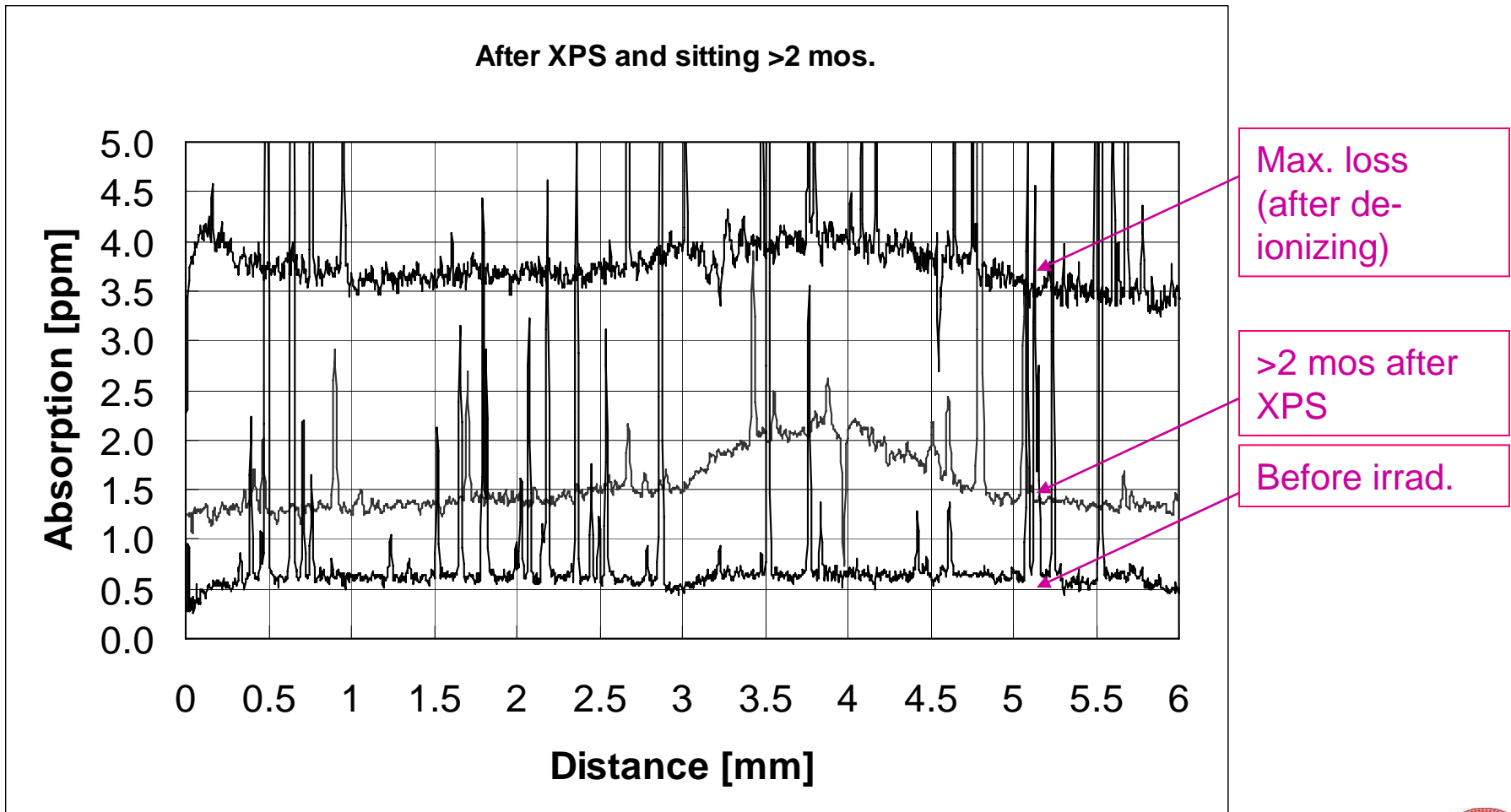


UV 'De-ionizer' as a Cleanser is Problematic

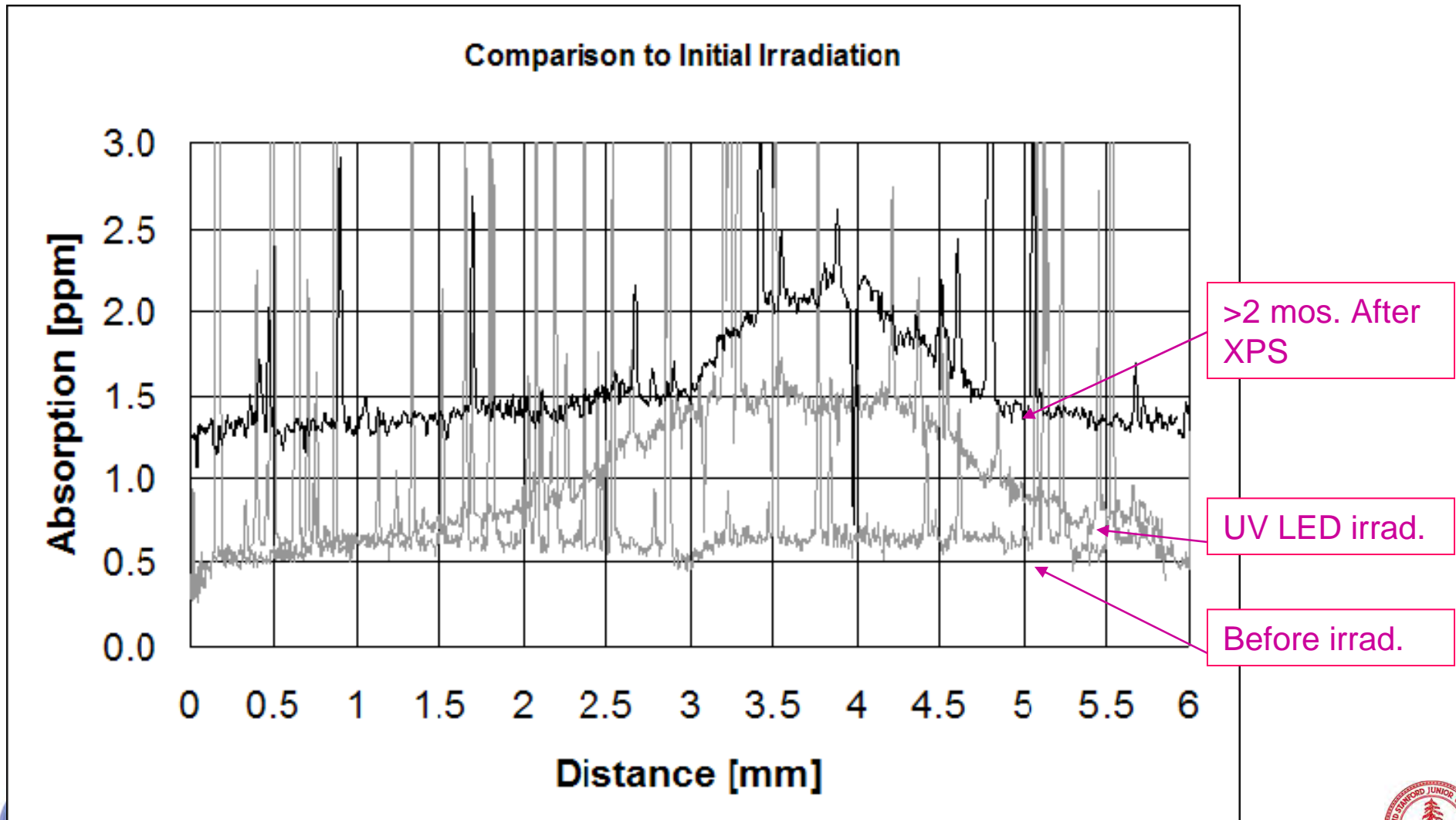
REO Sample #1, Deionizer Exposure in Air with Oxygen



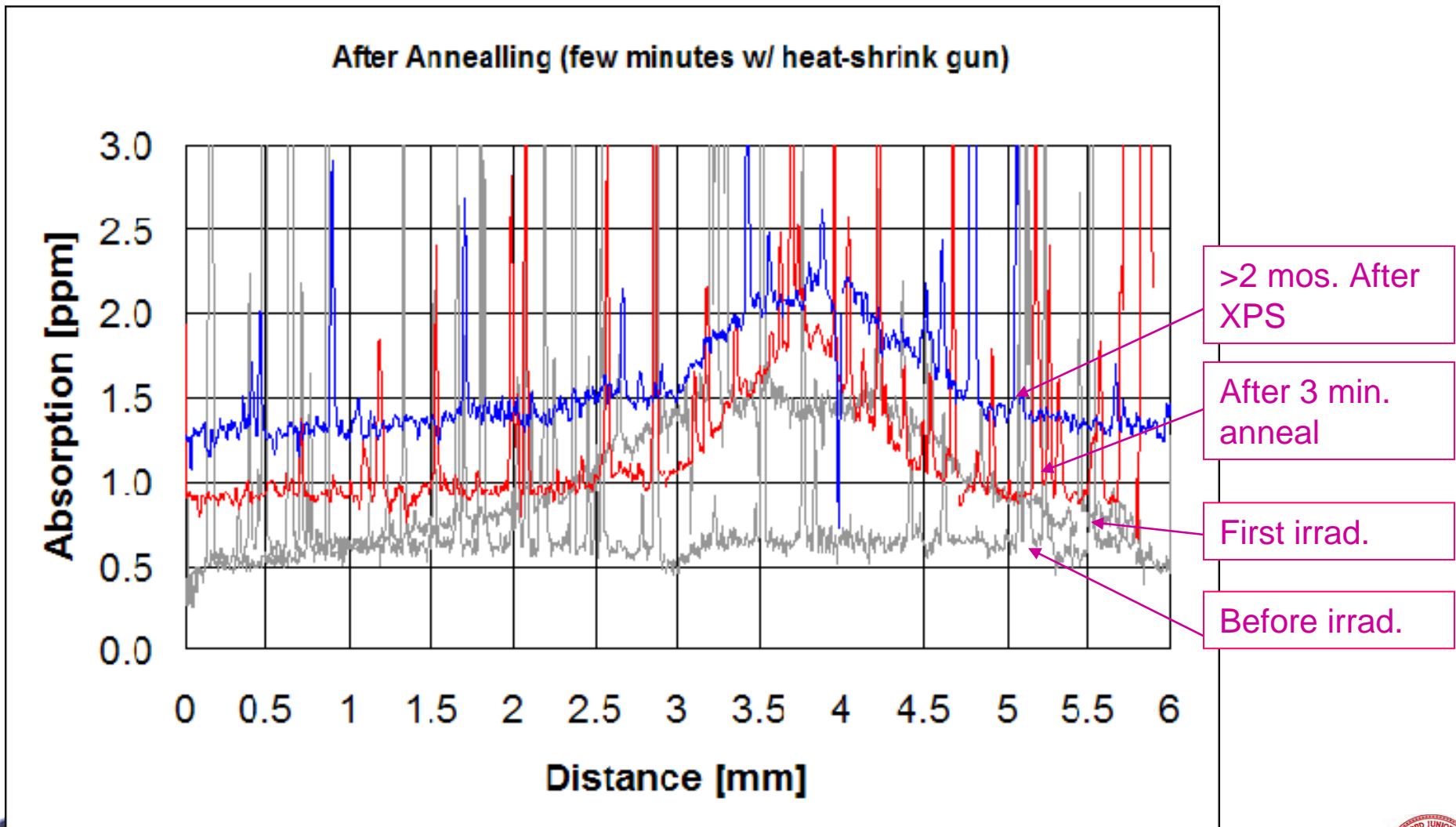
Time Heals the Wound?



Looking back



Effects of Dark Storage and Heat/Cool Cycles





Other Charging Management Techniques

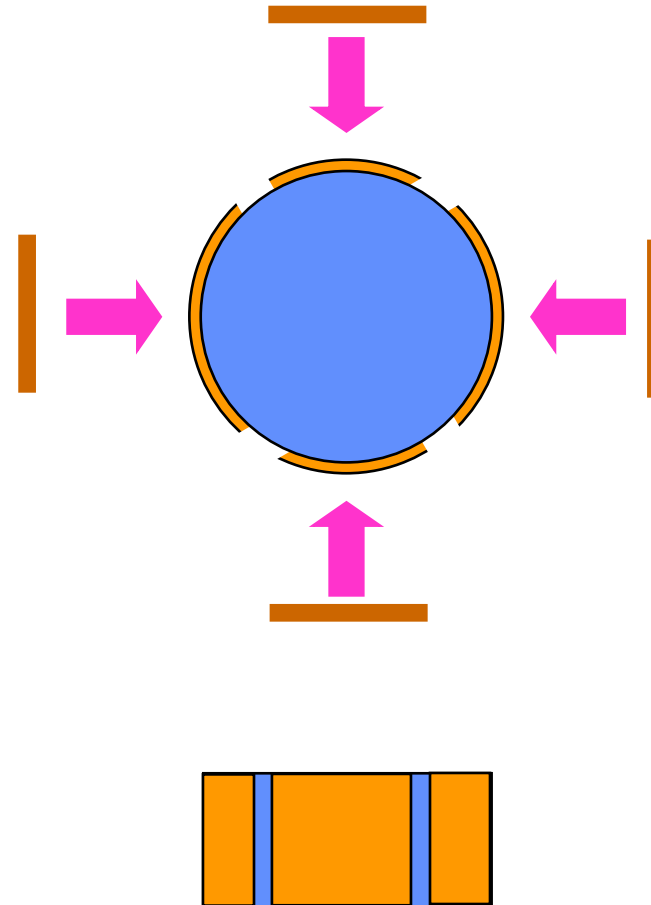
- **LIGO charge management is complicated by the dielectric core**
- **We are still committed to deliver a functional charge management solution to Advanced LIGO**
- **Devise indirect UV schemes for charge management**
 - **UV light does not illuminate the LIGO test mass material**
 - **UV light only produce electrons nearby LIGO test mass**
 - **External voltage (few volts) to steer electrons to equalize the potential**
 - **UV LED passed vacuum operation tests (ongoing for 3 months by far)**
- **Conductive coatings, wires, mass**
 - **Fluoride (CaF_2 , MgF_2 Often used in UV optics for lithography at 179~193 nm)**
 - **Patterned Au thin film on the barrel**
- **Ion and electron beam**
 - **Alternate ion beam and electron beam**
- **Neutralize the charges on test mass surface**

Au



Indirect Photoelectric Effect for Charge Management

- UV (violet) light not to directly illuminate the test mass
- Illuminate only segmented gold coating on barrel
- Allow electric charge to build up as a potential and field configuration to drive and deplete the charge on the test mass metal coating
- Then remove the charge from the metals



Conductive Coating and Test Mass

CaO:CaF₂

Influence of Oxygen on Electrical Properties of CaF₂ Crystals

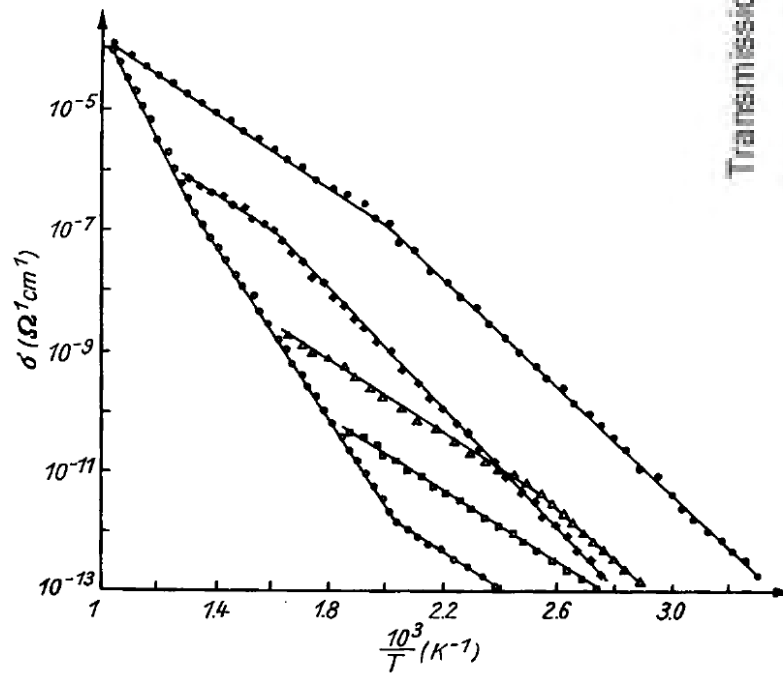
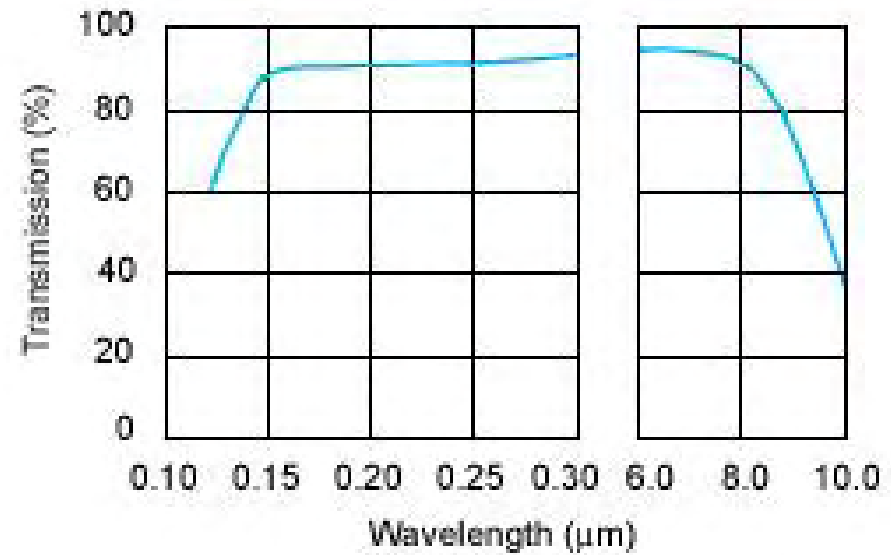


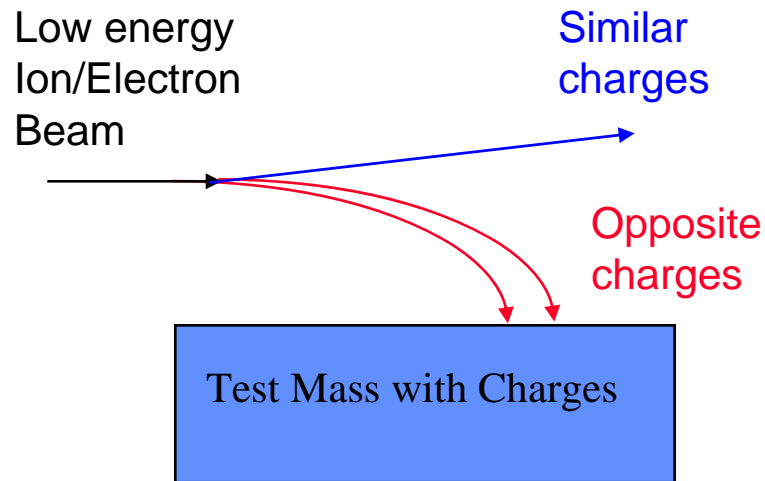
Fig. 1. Electrical conductivity of CaF₂ crystals. Curve 1: ○○○ raw material melted for 2 hours in products of Teflon pyrolysis, and dried previously for 48 hours; Curve 2: □□□ ibid, dried for 24 hours only; Curve 3: △△△ raw material melted for 2 hours in Teflon pyrolysis products; Curve 4: ■■■ CaF₂ crystal prepared at IKAN, Moscow, analyzed 0.01 mole % of oxygen; Curve 5: ●●● CaF₂ crystal with 0.1 mole % CaO



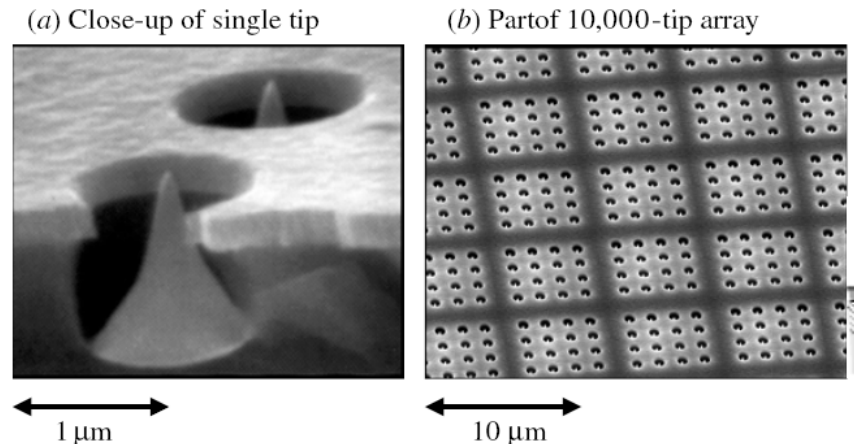
- CaF₂
- Transmission 170 nm – 8000 nm
- Ionic conductor
- Electrical conductivity improved by mixing in oxygen CaO



Ion and Electron Beam Charge Management



Spindt (SRI) Cathode Ion Sources



S Buchman¹, R L Byer, D Gill¹, N A Robertson and K-X Sun, “Charge neutralization in vacuum for non-conducting and isolated objects using directed low-energy electron and ion beams,” *Class. Quantum Grav.* **25** (2008) 035004

Ricardo suggested ionized argon gas by flow discharge

Summary

- **Charges on LIGO mirror Coating MUST be mitigated, especially for Advanced LIGO**
- **LIGO charge management present unique challenges beyond that from LISA**
- **UV light removes charges**
- **Systematic UV exposure effect measurement on LIGO and Adv LIGO samples**
- **UV effects on Ta₂O₃ based coating is small but real for bare dielectric coating**
- **Indirect UV illumination may be more effective and mostly safe for LIGO coatings**
- **Explore alternative charge management strategies**

